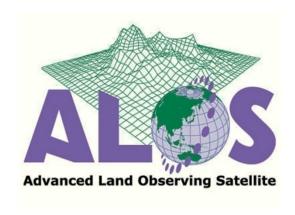


# Detection and Characterization of Ionospheric Effects in ALOS PALSAR data



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#### **Outline**



- Motivation
  - Project purpose
  - FR from Full-Pol data
    - Limitations and examples
  - FR from Dual-Pol data
    - Limitations, dependencies, and examples
- Background
  - lonosphere effects on Polarimetric SAR image calibration
  - Methods to estimate Faraday
     Rotation

- Evaluation
  - Comparison of methods
  - Comparison of different land classes
  - Full vs Dual Pol stats
  - Dual-pol attempts
    - Mountains
    - Oceans
- Conclusion
- Future



#### **Motivation**



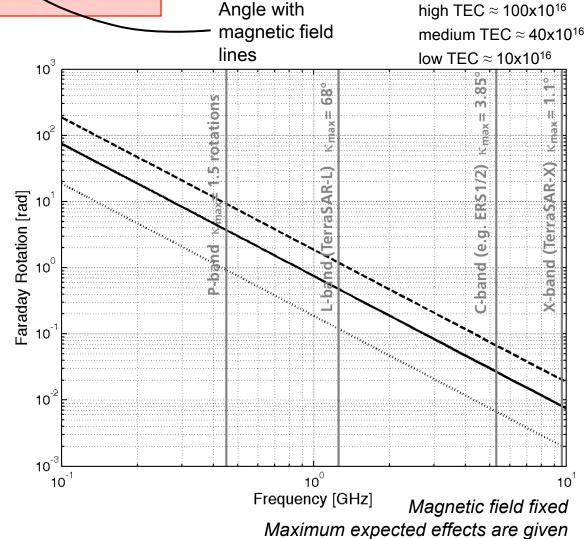
- ASF is working on an assessment of ionospheric influence on L-band SAR
  - Establish an operational monitoring/screening procedure to assist analysis of ionospheric effects and produce interesting data for ionospheric science
- Faraday rotation estimation is used as established reference technique
  - Requires full-polarimetric data sets
- A new method for ionospheric mapping from dual-pol data is presented and its performance is evaluated
  - May increase number of available data sets



# Faraday Rotation Effects on SAR 🛝



- Faraday Rotation:  $\Omega = \frac{K}{f^2} B \cos \theta \sec \chi \bullet TEC$
- Rotates energy from co-pol channels into cross-pol channels
  - Darker images, reduced signal-to-noise ratio, increased cross-talk, ...
  - Scattering matrix asymmetric
- More severe for L- and P-Band than for X- and C-Band
- Currently at solar low
  - Low TEC values dominate
  - Likely to continue for next 3 years
  - Events can still be dramatic



Look angle



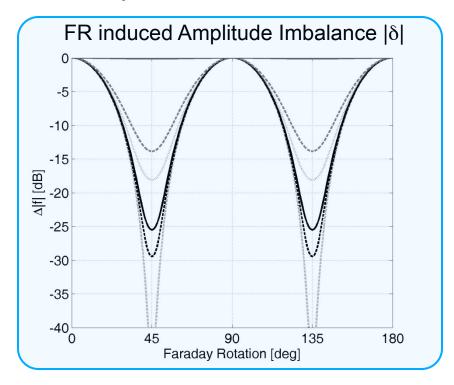
# Ionosphere Effects on Polarimetric SAR Image Calibration

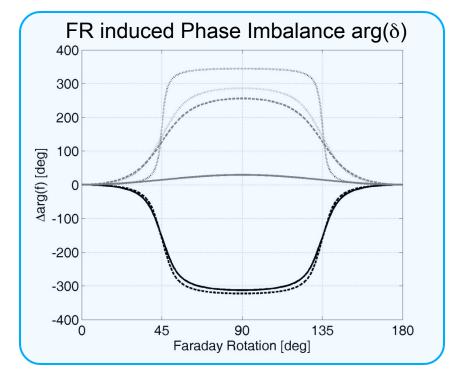


• For a full-polarimetric system:

$$\begin{bmatrix} M_{hh} & M_{hv} \\ M_{vh} & M_{vv} \end{bmatrix} = A e^{j\phi} \cdot \begin{bmatrix} 1 & \delta_1 \\ \delta_2 & f_1 \end{bmatrix} \cdot \begin{bmatrix} \cos \Omega & \sin \Omega \\ -\sin \Omega & \cos \Omega \end{bmatrix} \cdot \begin{bmatrix} S_{hh} & S_{hv} \\ S_{vh} & S_{vv} \end{bmatrix} \cdot \begin{bmatrix} \cos \Omega & \sin \Omega \\ -\sin \Omega & \cos \Omega \end{bmatrix} \cdot \begin{bmatrix} 1 & \delta_1 \\ \delta_2 & f_1 \end{bmatrix} + \begin{bmatrix} N_{hh} & N_{hv} \\ N_{vh} & N_{vv} \end{bmatrix}$$

Faraday rotation creates additional channel imbalance, and cross talk







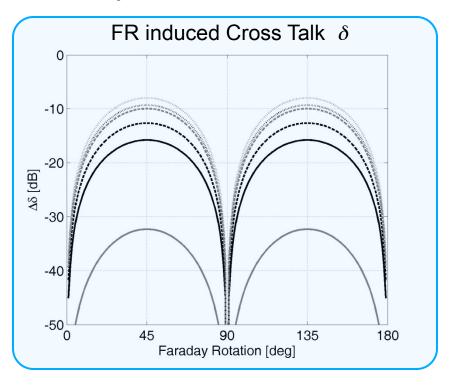
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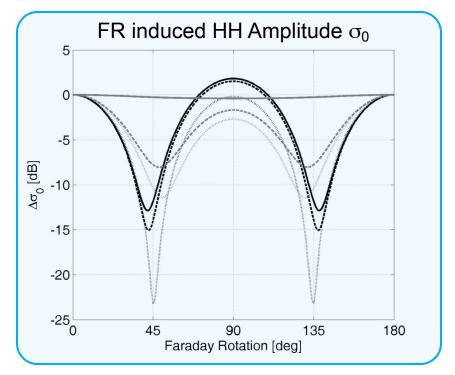


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Faraday rotation creates additional channel imbalance, and cross talk







### Faraday Rotation from Quad-Pol Data



#### Measured Scattering matrix of a sufficiently calibrated SAR system

$$\begin{bmatrix} M'_{hh} & M'_{vh} \\ M'_{hv} & M'_{vv} \end{bmatrix} = \begin{bmatrix} \cos \Omega & \sin \Omega \\ -\sin \Omega & \cos \Omega \end{bmatrix} \cdot \begin{bmatrix} S_{hh} & S_{vh} \\ S_{hv} & S_{vv} \end{bmatrix} \cdot \begin{bmatrix} \cos \Omega & \sin \Omega \\ -\sin \Omega & \cos \Omega \end{bmatrix}$$

Direct estimation from scattering matrix (Freeman, 2004):

$$\Omega = \frac{1}{2} \tan^{-1} \left[ \frac{\left( M'_{vh} - M'_{hv} \right)}{\left( M'_{vv} + M'_{vv} \right)} \right]$$

 $\Omega = \frac{1}{2} \tan^{-1} \left[ \frac{\left( M'_{vh} - M'_{hv} \right)}{\left( M'_{vv} + M'_{vv} \right)} \right]$  More robust version of estimator based on averaged second-order statistics published in (Freeman, 2004)

Estimation from circular basis (Bickel & Bates, 1965):

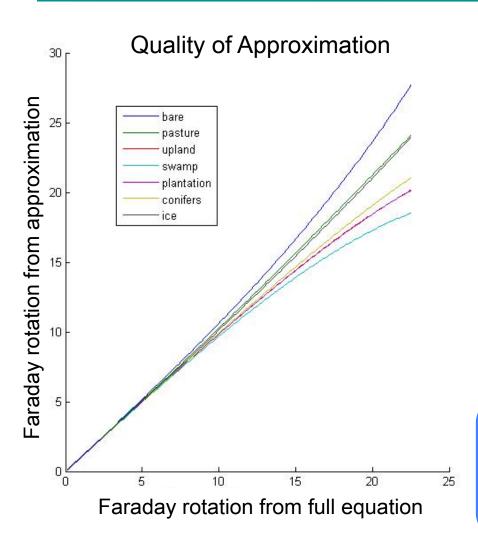
$$\begin{bmatrix} Z_{11} & Z_{12} \\ Z_{21} & Z_{22} \end{bmatrix} = \begin{bmatrix} 1 & j \\ j & 1 \end{bmatrix} \cdot \begin{bmatrix} M'_{hh} & M'_{vh} \\ M'_{hv} & M'_{vv} \end{bmatrix} \cdot \begin{bmatrix} 1 & j \\ j & 1 \end{bmatrix}$$

$$\Omega = \frac{1}{4} \arg \left( Z_{12} Z_{21}^* \right)$$



### Faraday Rotation Estimation





$$\Omega \approx \tan^{-1} \left( \frac{\left| \left\langle M_{hh} M_{hv}^* \right\rangle \right|}{\left| \left\langle M_{hh} M_{hh}^* \right\rangle \right|} \right)$$

where:

$$\frac{\left\langle S_{hh} S_{vv}^* \right\rangle}{\left\langle S_{hh} S_{hh}^* \right\rangle} = n = \left(\frac{\sigma_{vv}^0}{\sigma_{hh}^0} \cos(\phi_{hh-vv}) \rho_{hh-vv}\right)$$

*n* represents surface dependent scattering properties

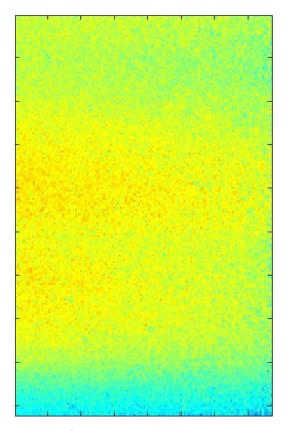
*n* can be derived in two ways:

- Calculated from full-pol data
- Extracted from scattering models

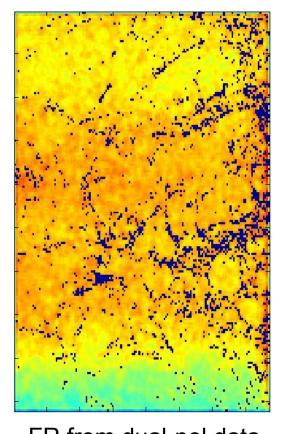


# Comparison of quad-pol and dual-pol methods assuming known *n*



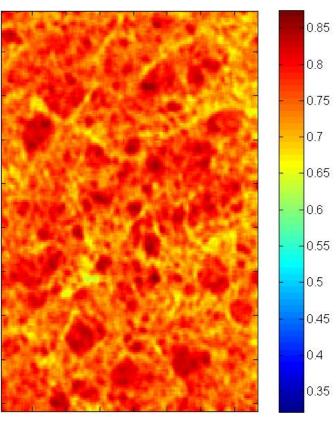


FR from Full-pol data



FR from dual-pol data

(Dark blue areas are masked out)



n calculated from full-pol data



### FR from quad-pol and dual-pol methods

# SAL S Advanced Land Observing Satellite

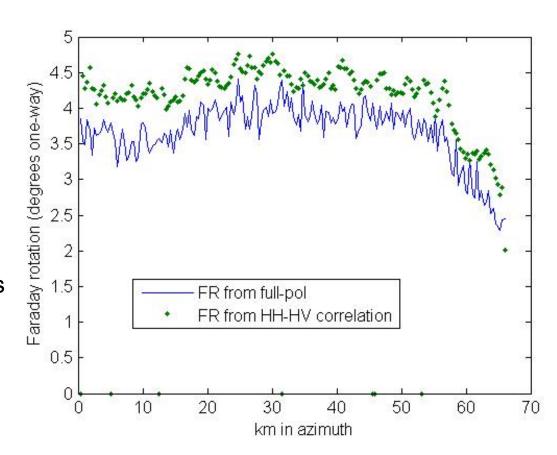
Assuming known *n* 

Transect along azimuth at midrange:

Strong agreement of applied methods

Estimates follow the same trend

Offset caused by approximations in FR equation (see slide 8)



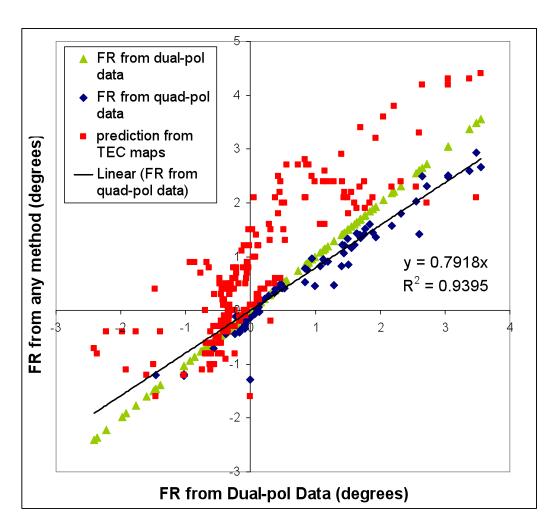


## FR from quad-pol and dual-pol methods

Fixed Value for *n* 



- FR from dual-pol data estimated with n fixed to n=0.7
  - n estimated empirically through minimization of estimation errors
- Comparison of dual-pol results to quad-pol estimates and predictions from global TEC maps
  - Dual-pol results are used as reference
- High correlation between dualpol and quad-pol estimates
  - Slight bias with increasing  $\Omega$





### FR from quad-pol and dual-pol methods

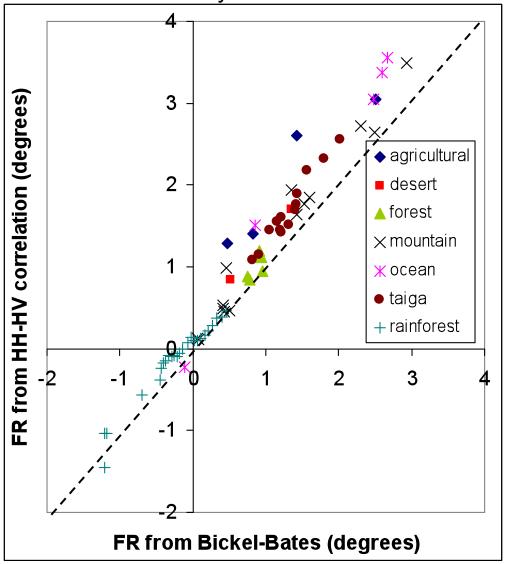
Fixed Value for *n* 



#### **Sensitivity to model errors**

- n varies with surface type
- Model fixes n to 0.7 → surface type dependent model errors
- No significant differences based on land type
  - Some classes are not wellsampled
  - Need more full-pol data to compare methods

#### Sensitivity to model errors

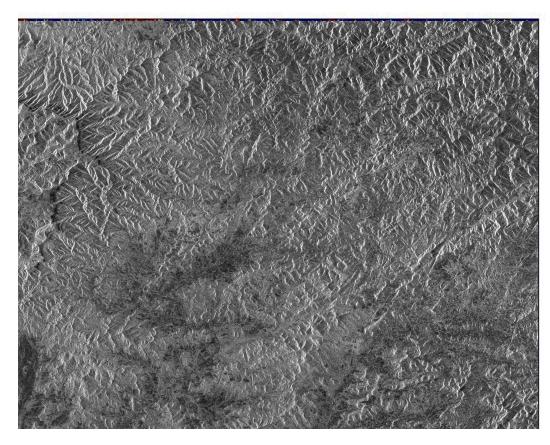




# Influence of Scattering Asymmetry



- In contrast to full-pol based FR estimation using Bickel&Bates methods, dual-pol method is sensitive to scattering asymmetry
- Slopes in azimuth direction cause scattering asymmetry
- Adds variation to the measurement
- If averaged over large areas, bias averages out
  - no apparent bias



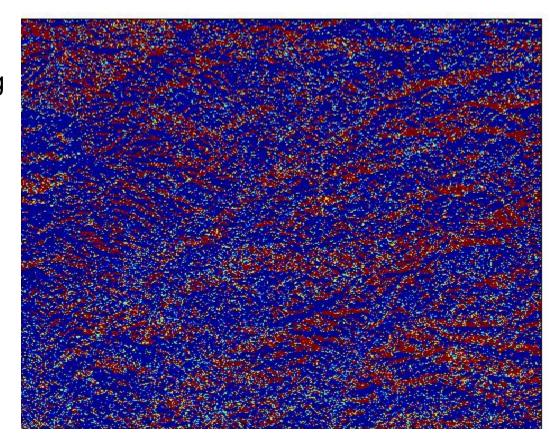
Normalized cross-correlation coefficient for dualpol scene ALPSRP073340230



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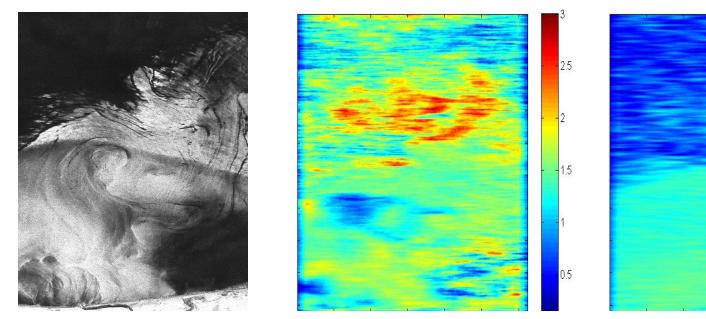


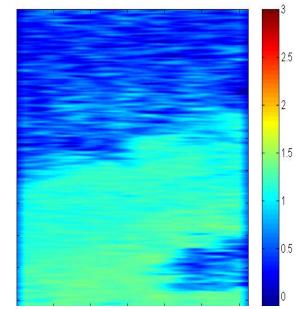
Normalized cross-correlation coefficient for dualpol scene ALPSRP073340230



# Influence of Signal-to-Noise Variations 🛕







Original (left), FR from cross-channel (mid) and Bickel-Bates (right) for a scene with water and land.

Note that Bickel-Bates underestimates water (due to low signal to noise) Dual-pol estimate doesn't show the same behavior → under investigation

ALPSRP063051200



#### Conclusions



- Ionospheric effects influence the calibration quality of SAR and PolSAR data
- Significance of ionospheric influence is currently largely unknown due to missing statistical parameters
- ASF is preparing a system for continuous ionospheric mapping from PALSAR data to provide these statistical parameters
- Main estimation methods is Faraday rotation
- HH/HV Correlation applies to dual-pol data and increases number of available data sets
- First analyses show that Bickel & Bates as well as HH/HV correlation methods produce reliable results
- Influence of surface type variation appears to be limited for low FR



#### **Future Work**



- Finalizing of performance analysis of HH/HV Cross correlation
- Finalizing operational implementations of mapping techniques
- Processing full-pol and dual-pol data in the data pool over high-latitude areas to assess effects of more turbulent and less predictable polar ionosphere
- Additionally, processing of all new datasets that are transferred to the data pool
- Implementation of additional estimation methods (e.g. incoherent autofocus techniques) will be prepared

Thanks for your attention!!

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